

Ultracold mixtures of metastable He and Rb: scattering lengths from *ab initio* calculations and thermalization measurements

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We have investigated the ultracold interspecies scattering properties of metastable triplet He and Rb. We performed state-of-the-art *ab initio* calculations of the relevant interaction potential, and measured the interspecies elastic cross section for an ultracold mixture of metastable triplet ⁴He and ⁸⁷Rb in a quadrupole magnetic trap at a temperature of 0.5 mK. Our combined theoretical and experimental study gives an interspecies scattering length $a_{4+87} = +17^{+1}_{-4} a_0$, which prior to this work was unknown. More general, our work shows the possibility of obtaining accurate scattering lengths using *ab initio* calculations for a system containing a heavy, many-electron atom, such as Rb.

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I. INTRODUCTION

Ultracold mixtures of different atomic species offer a number of advantages over single species experiments. For instance, these mixtures are the starting point to obtain a dense sample of ultracold heteronuclear molecules, which (in contrast to homonuclear molecules) can have long-range and anisotropic interactions, resulting in rich new physics with many novel properties [1]. Ultracold mixtures can also feature very interesting few- and many-body phenomena, for which the mass ratio between the two atomic species can play a crucial role (see e.g. [2]). A prominent example is the observation of more than two successive Efimov loss features to test the scaling laws of the Efimov trimer spectrum, which experimentally requires an extremely large mass ratio, and for which first results have been obtained in ultracold mixtures of ⁶Li+¹³³Cs [3].

Most experiments on ultracold mixtures involve two alkali-metal species, while recently also mixtures of alkali-metal and alkaline-earth(-like) atoms became available [4]. Here we are considering a different type of mixture, namely of an alkali-metal atom and helium, in the metastable 2^3S_1 triplet state (denoted as He*, radiative lifetime of about 8000 s), for which quantum degeneracy has been realized for both fermionic ³He* and bosonic ⁴He* isotopes [5]. The application of He* in ultracold mixtures increases the range of possible mass ratios by a factor of two compared to the commonly used ⁶Li.

The feasibility of an ultracold or quantum degener-

ate mixture depends strongly on the collisional properties and stability, which in turn is given by the intraspecies and interspecies interaction potentials. Scattering between He* (total electron spin $s=1$) and an alkali-metal atom ($s=1/2$) in the electronic ground state is described by a doublet $^2\Sigma^+$ and a quartet $^4\Sigma^+$ molecular potential. Here we focus on the $^4\Sigma^+$ potential, for which Penning ionization is suppressed due to spin-conservation [6] and which fully describes a mixture in which both atoms are either in the lower or upper spin-stretched states. These spin-mixtures are most favorable for sympathetic and evaporative cooling towards quantum degenerate mixtures. Precise knowledge of those potentials is completely lacking, due to the absence of spectroscopic data. Therefore one has to rely on *ab initio* calculations, for which however the predicted power in terms of accurate scattering lengths is generally considered to be limited, except for few-electron systems like He*+He* [7].

In this article we present state-of-the-art *ab initio* quantum chemistry calculations of the $^4\Sigma^+$ potential and the quartet scattering lengths for the He*+Rb system. In parallel, we have experimentally determined the quartet scattering length for ⁴He*+⁸⁷Rb by measuring the interspecies elastic cross section for an ultracold mixture in a quadrupole magnetic trap. Our combined theoretical and experimental work gives tight bounds on the interspecies scattering lengths, which provides crucial knowledge for the realization of quantum degenerate He*+Rb mixtures.

This paper is organized as follows. In Sec. II we present the *ab initio* calculations. In Sec. III we describe the experiment, including a theoretical description of interspecies thermalization measurements in a quadrupole magnetic trap (Sec. III A), the experimental setup (Sec. III B), the two-species magneto-optical trap (Sec. III C) and quadrupole magnetic trap

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(Sec. IIID), and the determination of the scattering length (Sec. IIIE). In Sec. IV we compare the theoretical and experimental results and conclude. Finally, in Sec. V we give some future prospects.

II. AB INITIO CALCULATIONS

The calculations of the $^4\Sigma^+$ potential have been performed using the all-electron restricted open-shell coupled cluster singles and doubles with noniterative triples corrections [CCSD(T)] method [8], implemented in the MOLPRO package [9], and using the Douglas-Kroll-Hess Hamiltonian to take into account relativistic effects [10]. These calculations are challenging for a system like He*Rb, since the molecular states are submerged in the continuum of ionized states of HeRb⁺, which might in principle lead to a variational collapse to lower lying states already during the Hartree-Fock (HF) optimization [11]. To circumvent this, we have constructed starting orbitals from appropriate orbitals of the isolated atoms and during the optimization we have kept the occupancies of orbitals fixed.

For He*Rb the available standard electronic gaussian basis sets are not appropriate, in particular because the basis sets for He are optimized to recover the ground-state energy. Therefore we have optimized our own basis set, suitable for He*. For Rb we have used the ANO-RCC basis set [12], to which we have added one *g*- and two *h*-type orbitals optimized to the atomic energies. To better account for the dispersion interaction we have augmented both basis sets using two sets of even-tempered functions per function type (generated with the MOLPRO package). The convergence of the counterpoise-corrected interaction energies [13] is carefully analyzed both in terms of the number of augmented functions added, as well as the highest angular momentum function in the basis set. By removing the *h* basis functions we have found that the interaction energy changes by less than 1 cm⁻¹.

The coupled-clusters equations are divergent for internuclear distances smaller than $r = 8 a_0$, for which the interaction energy is approximately -200 cm^{-1} and the inner turning point is not yet reached. Still, we were able to converge the HF reference state down to $r = 5.5 a_0$, from which we can exclude the possibility of crossings of the potential energy curve with other states. To extrapolate the potential towards shorter distances we have made use of the fact that the contribution of correlation energy to the interaction potential, which is by far dominated by the dispersion energy, varies exponentially near the inner turning point [14] and added the extrapolated values to the HF interaction energy (see Appendix A 1).

The long-range van der Waals coefficients have been recently calculated by Zhang *et al.* [15], however, with an uncertainty of 1-5% in C_6 and 1-10% in C_8 and C_{10} , which is too large for our purpose. We therefore have calculated the dipolar- and quadrupole dynamic polarizabilities of He* at imaginary frequencies and integrated

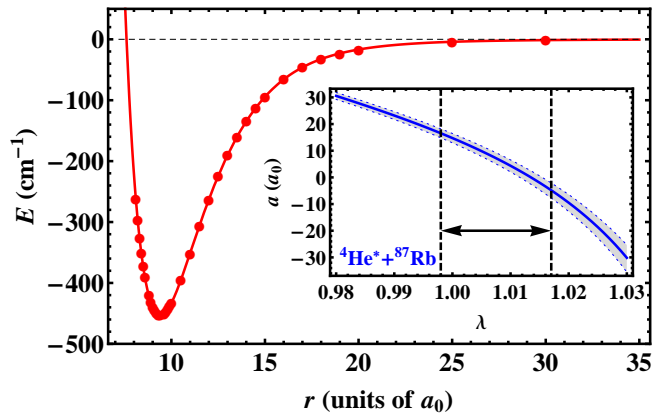


FIG. 1. (Color online) Results of the *ab initio* calculations on the $^4\Sigma^+$ potential of He*Rb (red circles) and the MLR fit (red solid line). The inset shows the quartet scattering length for $^4\text{He}^* + ^{87}\text{Rb}$ as function of the scaling parameter λ (see text), where the shaded area (bounded by the blue dotted lines) represents the uncertainties in the long-range coefficients. The dashed vertical lines and the arrow represent the uncertainty in the *ab initio* calculations, corresponding to a range of $\lambda=0.998\text{--}1.017$.

TABLE I. Parameter values of the MLR potential.

parameter	value	parameter	value
D_e	452.71 cm^{-1}	ϕ_0	-1.8284
r_e	$9.4079 a_0$	ϕ_1	0.48678
C_6	$8.4673 \times 10^8 \text{ cm}^{-1} a_0^6$	ϕ_2	-0.065081
C_8	$8.0108 \times 10^{10} \text{ cm}^{-1} a_0^8$	ϕ_3	-0.30087
C_{10}	$9.4242 \times 10^{12} \text{ cm}^{-1} a_0^{10}$	ϕ_4	-1.5195

them with recently tabularized dynamic polarizabilities of Rb. With the present calculations the error in the C_6 coefficient for Rb₂ system is estimated as 0.5%, while the C_6 and C_8 coefficients of He₂* reproduce the reference data [16] to better than 0.1%. For He*Rb we obtain $C_6 = 8.47(2) \times 10^8 \text{ cm}^{-1} a_0^6$, with an uncertainty of 0.25%. Using the single-pole approximation to the quadrupole dynamic polarizability derived by Porsev *et al.* [17], we have obtained also $C_8 = 8.01(4) \times 10^8 \text{ cm}^{-1} a_0^8$, which has an accuracy of 0.56%. Finally we include the C_{10} coefficient from Zhang *et al.* [15] to the long-range part of our potential.

The potential energy curve obtained from the *ab initio* calculations is shown in Fig. 1. The data points are fitted with the Morse/Long-Range (MLR) potential proposed by LeRoy *et al.* [18], which has the form:

$$V(r) = D_e \left[1 - \frac{u_{\text{LR}}(r)}{u_{\text{LR}}(r_e)} \exp(-\phi(r)y_p(r)) \right]^2 - D_e, \quad (1)$$

where D_e is the well depth of the potential, r_e the equilibrium distance, $u_{\text{LR}}(r) = C_6 r^{-6} + C_8 r^{-8} + C_{10} r^{-10}$, $y_k(r) = (r^k - r_e^k) / (r^k + r_e^k)$ and $\phi(r) =$

$[1 - y_p(r)] \sum_{j=0}^4 \phi_j y_q(r) + y_p(r) \phi_\infty$, where $p = 5$ and $q = 4$. The free parameters in the potential, determined by fitting, are ϕ_j ($j = 0, \dots, 4$), while D_e , r_e and $\phi_\infty = \log[2D_e/u_{\text{LR}}(r_e)]$ are directly obtained from the *ab initio* calculations. The resulting parameter values of the MLR potential are given in Table I. Note that the statistical error introduced by the analytical fit is much smaller than the systematic uncertainty in the *ab initio* calculations. The MLR potential is particularly convenient for the analysis of the scattering length in the case that the long-range part of the potential is known very accurately, but the short-range potential has a larger uncertainty. Hence we can conveniently parametrize the scattering length by introducing a λ scaling parameter such that $D_e \rightarrow \lambda D_e$.

To predict quantitatively the scattering length it is crucial to explore possible errors in the *ab initio* calculations. We therefore have used higher-order coupled cluster methods, using the MRCC code [19], to estimate the uncertainty in the potential energy curve. We have found that the systematic error that can be attributed to the incompleteness of CCSD(T) is approximately $+4.3 \text{ cm}^{-1}$ (see Appendix A2). We have also compared the potential depths for the homonuclear He^* and Rb dimers obtained with the same method used in this work, and the reference D_e parameters for He_2^* [7] and Rb_2 [20]. The fractions $\delta(X_2) = [D_e^{\text{ref}}(X_2) - D_e^{\text{calc}}(X_2)] / D_e^{\text{ref}}(X_2)$ are respectively 5.4×10^{-3} and 4.1×10^{-2} , hence for the heteronuclear system we can estimate the error as $\sqrt{\delta(\text{He}_2^*)\delta(\text{Rb}_2)} = 0.015$, which translates into $+6.8 \text{ cm}^{-1}$. Finally, the long-range CCSD(T) potential curve can be tested by comparing it with the $V_{\text{LR}}(r) = -u_{\text{LR}}(r)$ expansion of the interaction energy. For distances between $25 a_0$ and $40 a_0$, where the potential energy is dominated by the interaction of multipoles, $V_{\text{LR}}(r)$ is systematically larger by about 1.2% than the CCSD(T) potential, which corresponds to a difference in D_e of $+5.4 \text{ cm}^{-1}$. All estimations on a possible error in D_e give a systematically positive shift. Therefore we conservatively assume the error bound between -1 cm^{-1} (uncertainty of basis set) and $+7.8 \text{ cm}^{-1}$ (uncertainty of basis set and most conservative estimate of the CCSD(T) uncertainty), which translates to a scaling parameter range of $\lambda = 0.998 - 1.017$.

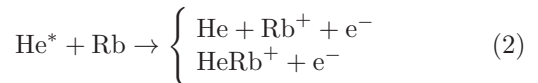
From the potential energy curve we calculate the scattering lengths for all the four isotope combinations, for which the results are given in Table II. The inset of Fig. 1 shows the scattering length for $^4\text{He}^* + ^{87}\text{Rb}$ as function of λ . We find that the scattering lengths for all isotope combinations are small, i.e. $|a| \leq 21 a_0$. The small difference between ^{85}Rb and ^{87}Rb for a given He^* isotope is due to the small difference in reduced mass. In contrast, the small difference between $^4\text{He}^* + \text{Rb}$ and $^3\text{He}^* + \text{Rb}$, for which the reduced mass is very different, is completely accidental. For instance, the $^4\text{He}^* + \text{Rb}$ potential supports 15 bound states, compared to 13 for the $^3\text{He}^* + \text{Rb}$ potential.

TABLE II. The quartet scattering lengths of $\text{He}^* + \text{Rb}$ in units of Bohr radius a_0 , obtained from the *ab initio* calculations (theory) and thermalization measurements (experiment). For the theory we give the values connected to $\lambda=1$ and the bounds corresponding to $\lambda=[1.017; 0.998]$. The experimentally obtained interspecies elastic cross section give rise to two possible values of the scattering length (see Sec. III E), where the error bars corresponds to one standard deviation.

isotopes	3+85	3+87	4+85	4+87
theory	+6 [-19; +8]	+5 [-21; +7]	+18 [-5; +20]	+16 [-8; +18]
experiment	-29_{-5}^{+5} or $+17_{-4}^{+4}$			

III. EXPERIMENT

In the following we discuss the thermalization measurements, including the experimental setup and our strategy to obtain an ultracold mixture in a quadrupole magnetic trap (QMT), in which both species are in their fully stretched magnetic substate, i.e. $^4\text{He}^*$ in the $J=1$, $m_J=1$ state and ^{87}Rb in the $F=2$, $m_F=2$ state. The reason for choosing this *doubly spin-stretched mixture* is that interspecies Penning and associative ionization processes (which we both will refer to as PI), i.e.



are expected to be suppressed because of spin-conservation [5]. An upper limit for the loss rate coefficient of $5 \times 10^{-12} \text{ cm}^3 \text{ s}^{-1}$ at 0.2 mK has been experimentally obtained by measuring the ion production rate [6]. For other spin-mixtures large loss rate coefficients on the order of $10^{-10} \text{ cm}^3 \text{ s}^{-1}$ are expected [5]. Simultaneous laser cooling and trapping of $^4\text{He}^*$ and ^{87}Rb has already been demonstrated by the Truscott group [6, 21].

A. Theoretical description of interspecies thermalization in QMT

Interspecies thermalization of ultracold mixtures has been described in detail in many papers (see e.g. [22]), although mostly for a Ioffe-Pritchard type of magnetic trap or optical dipole traps, i.e. a harmonic trapping potential. Here we consider thermalization for a QMT, i.e. a linear trapping potential, which requires the inclusion of Majorana heating.

The time evolution of the temperature difference $T_1 - T_2$ in a two-species mixture is described by

$$\frac{d}{dt} (T_1 - T_2) = -\gamma_{\text{th}} (T_1 - T_2), \quad (3)$$

with thermalization rate $\gamma_{\text{th}} = \gamma_{\text{coll}} \xi / 2.7$, where γ_{coll} is the collision rate and $\xi = 4m_1 m_2 / (m_1 + m_2)^2$. For equal

mass systems 2.7 collisions are required for thermalization [23], which can be generalized to $2.7/\xi$ for non-equal masses [22]. The collision rate is given by $\gamma_{\text{coll}} = \sigma \langle v \rangle \langle n \rangle$, with the interspecies elastic cross section σ , the mean velocity $\langle v \rangle = \sqrt{8k_B/\pi (T_1/m_1 + T_2/m_2)}$ and the mean density $\langle n \rangle = (1/N_1 + 1/N_2) \int n_1(\vec{r})n_2(\vec{r})d\vec{r}$. The temperature dependence of σ will be discussed in Appendix B.

In a QMT the density distribution (assuming an infinitely deep trap) is given by

$$n(x, y, z) = n_0 \exp \left[-\frac{\mu\alpha\sqrt{x^2 + 4y^2 + z^2} - mgz}{k_B T} \right], \quad (4)$$

where in our case the axial direction of the coils (y -axis) is in the horizontal plane, μ is the magnetic moment, α is the magnetic field gradient along the weak (radial) axis, g is the gravitational acceleration and the peak density

$$n_0 = \frac{N}{4\pi} \left(\frac{\mu\alpha}{k_B T} \right)^3 \left[1 - \left(\frac{mg}{\mu\alpha} \right)^2 \right]^2. \quad (5)$$

In our case we can safely neglect the effect of gravity, as $mg/\mu\alpha$ is small (0.13 for ^{87}Rb , 0.003 for $^4\text{He}^*$), and the reduction of the overlap $\langle n \rangle$ caused by the gravitational sag of the ^{87}Rb distribution is less than 3%. In this approximation, $\int n_1(\vec{r})n_2(\vec{r})d\vec{r} = (\alpha^3 N_1 N_2 / 4\pi k_B^3) (T_1/\mu_1 + T_2/\mu_2)^{-3}$ and the thermalization rate is given by

$$\gamma_{\text{th}} = \frac{\sigma \xi \alpha^3 (N_1 + N_2)}{2.7 \sqrt{2} \pi^{3/2} k_B^{5/2}} \frac{\sqrt{\frac{T_1}{m_1} + \frac{T_2}{m_2}}}{\left(\frac{T_1}{\mu_1} + \frac{T_2}{\mu_2} \right)^3}. \quad (6)$$

In case $N_1 \gg N_2$, γ_{th} does not depend on N_2 .

In addition we have to include the Majorana effect, i.e. nonadiabatic spin flips to untrapped states at the magnetic field zero at the center of the QMT, which leads to both losses and heating, and therefore limits evaporative cooling [24], but also interspecies thermalization and sympathetic cooling. The Majorana heating rate is described by [25]

$$\frac{dT}{dt} = \frac{\gamma_{\text{Maj}}}{2T}, \quad (7)$$

where $\gamma_{\text{Maj}} = (8/9)\chi(\hbar/m)(2\mu\alpha/k_B)^2$, and χ is a dimensionless factor. The solution of Eq. 7 is given by $T(t) = \sqrt{T_0^2 + \gamma_{\text{Maj}} t}$, where T_0 is the initial temperature.

The combined effect of interspecies thermalization and Majorana heating is then described by

$$\begin{aligned} \frac{d}{dt} (T_1 - T_2) = & -\gamma_{\text{th}} (T_1 - T_2) + \frac{\gamma_{\text{Maj},1}}{2T_1} - \frac{\gamma_{\text{Maj},2}}{2T_2} \\ & + \left[\frac{dT_1}{dt} \right]_{\text{ev}} - \left[\frac{dT_2}{dt} \right]_{\text{ev}}, \end{aligned} \quad (8)$$

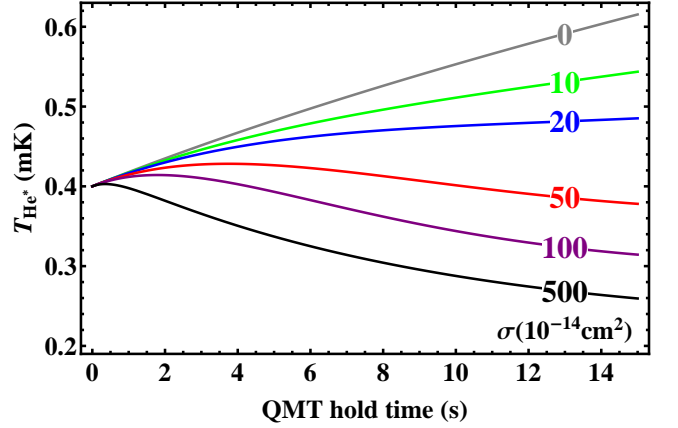


FIG. 2. (Color online) Calculated thermalization curves for different interspecies elastic cross sections σ , with $\alpha = 120$ G/cm, $\chi=0.1$, initial temperature $T_{\text{He}^*} = 0.4$ mK, and atom numbers $N_{\text{He}^*}^0 = 5 \times 10^6$ and $N_{\text{Rb}}^0 = 2 \times 10^8$, where for $^4\text{He}^*$ and ^{87}Rb we have included a one-body loss rate of $(30 \text{ s})^{-1}$ and $(15 \text{ s})^{-1}$, respectively. For the ^{87}Rb temperature we assume $T_{\text{Rb}} = (T_0 - T_f)e^{-\beta t} + T_f$, with $T_0=0.4$ mK, $T_f=0.2$ mK and $\beta = 0.1 \text{ s}^{-1}$.

where the last two terms include the effect of evaporative cooling. For our experimental parameters we can neglect Majorana heating for ^{87}Rb (γ_{Maj} is a factor 87 smaller than that of $^4\text{He}^*$). As we will show below, the time evolution of the ^{87}Rb temperature is due to plain evaporation, i.e. $dT_{\text{Rb}}/dt = [dT_{\text{Rb}}/dt]_{\text{ev}}$, whereas for $^4\text{He}^*$ the trap depth is too large for evaporative cooling, i.e. $[dT_{\text{He}^*}/dt]_{\text{ev}} = 0$. This all means that for our situation we can effectively simplify Eq. 8 to:

$$\frac{d}{dt} T_{\text{He}^*} = -\gamma_{\text{th}} (T_{\text{He}^*} - T_{\text{Rb}}) + \frac{\gamma_{\text{Maj,He}^*}}{2T_{\text{He}^*}}, \quad (9)$$

where it is important to note that γ_{th} depends on T_{He^*} , T_{Rb} and N_{Rb} , which all change during the hold time in the QMT. The solution of Eq. 9 for different values of σ is shown in Fig. 2.

B. Experimental setup

A schematic of the setup is shown in Fig. 3. We have added a 2D-MOT for Rb on one of the viewports of the stainless steel vacuum chamber of an existing He^* setup [26]. We use a liquid-nitrogen cooled dc-discharge source to produce a ^4He beam with a $^4\text{He}^*$ fraction of 10^{-4} [27]. The $^4\text{He}^*$ beam is collimated with a total power of about 500 mW, slowed in a 2.5 m long Zeeman slower and loaded into the 3D-MOT. An *in vacuo* shutter is opened only during the loading time of the $^4\text{He}^*$ 3D-MOT. Without the shutter the lifetime of $^4\text{He}^*$ and ^{87}Rb atoms in the QMT is limited to less than two seconds. The pressure in the main vacuum chamber is 1×10^{-10} mbar.

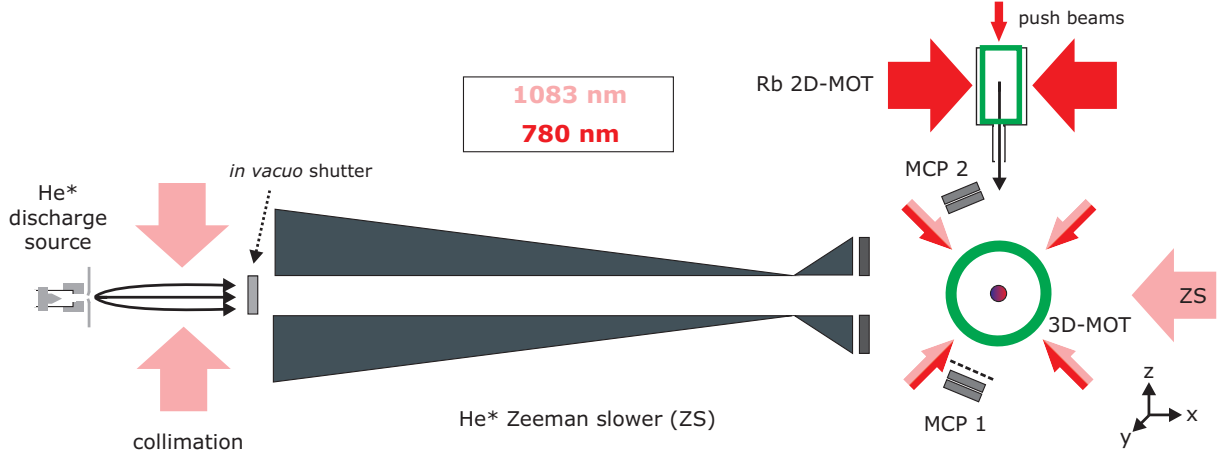


FIG. 3. (Color online) A schematic overview of the experimental setup and laser beams. The vertical direction is along the z -axis. Note that the pair of laser beams along the y -axis of the collimation section, the 2D-MOT and the 3D-MOT are not shown. Optical pumping and absorption imaging beams are along the y -axis (ZS: Zeeman slower; MCP: micro-channel plate).

For the detection of $^4\text{He}^*$ two micro-channel plate (MCP) detectors are placed at a distance of 106 mm from the trap center, both under an angle of 22° with respect to the direction of gravity. MCP 1 is behind a grounded grid and detects $^4\text{He}^*$ atoms after release from the trap, resulting in a time-of-flight (TOF) signal that contains information about the atom number and the temperature. MCP 2 is not shielded and therefore also collects all ions (He^+ , Rb^+) produced via PI. We have calibrated the MCP signals using the saturated fluorescence method, collecting transient fluorescence from a retro-reflected high-power resonant beam [28]. For ^{87}Rb we use standard absorption imaging (along y -direction) to obtain the atom number and temperature.

One pair of water-cooled coils provides the magnetic field gradient for both the 3D-MOT and QMT, which has a gradient along the weak axis $\alpha = \partial B/\partial x = \partial B/\partial z = \frac{1}{2}\partial B/\partial y$ of 0.6 (G/cm)/A . The axial direction of the coils (y) is in the horizontal plane. The laser beams for optical pumping and absorption imaging are along the y -direction.

C. Two species magneto-optical trap

The three retroreflected 1-inch laser beams of the 3D-MOT are derived from single mode optical fibers, in which both wavelengths for laser cooling of $^4\text{He}^*$ and ^{87}Rb , 1083 nm and 780 nm, respectively, are coupled together using dichroic mirrors. In this way, the 3D-MOT laser beams of the two species are automatically overlapped. To create the proper circular polarization for both wavelengths, zero-order quarter wave plates at 920 nm are used. For ^{87}Rb a detuning of -15 MHz with respect to the $F = 2 \rightarrow F' = 3$ transition is used, and a total power of the three laser beams of $\sim 40 \text{ mW}$. For $^4\text{He}^*$ we use a large detuning of -32 MHz (corresponding to 20 linewidths) to reduce the light-assisted

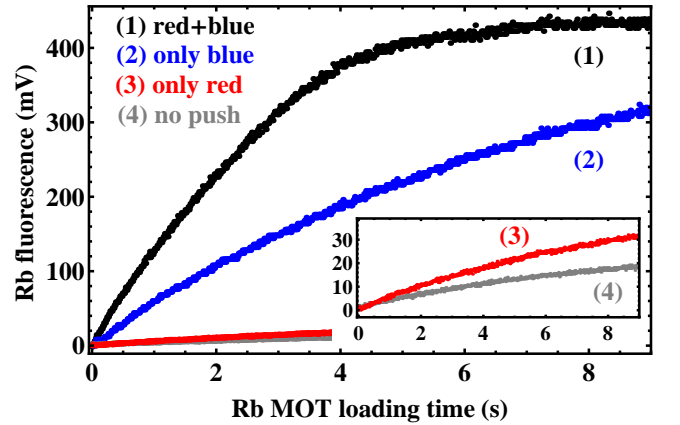


FIG. 4. (Color online) ^{87}Rb fluorescence during the loading of the ^{87}Rb 3D-MOT, showing different configurations of the 2D-MOT: (1) both red- and blue-detuned push beams (black), (2) only blue-detuned push beam (blue), (3) only red-detuned push beam (red), and (4) no push beams (gray). A fluorescence signal of 400 mV corresponds to about 1×10^9 atoms.

intraspecies PI loss in the 3D-MOT, and a total laser beam power of $\sim 30 \text{ mW}$. The magnetic field gradient α is 12 G/cm . For ^{87}Rb an additional repumper beam on the $F = 1 \rightarrow F' = 2$ transition is added. For $^4\text{He}^*$ no repumper is needed because of the absence of hyperfine structure.

^{87}Rb is loaded from a 2D-MOT with a two-color push beam scheme, in which a red-detuned push beam pushes the atoms that leave the 2D-MOT in the wrong direction back towards the 3D-MOT, while a blue-detuned push beam guides the atoms through the differential pumping tube [29]. The two retroreflected cooling beams are circular with a diameter of two inches and a total power of the two laser beams of $\sim 100 \text{ mW}$. The detuning is

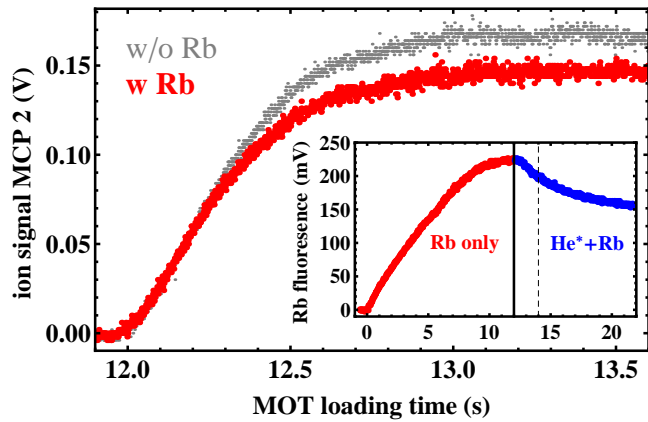


FIG. 5. (Color online) Two species MOT loading. Ion signal of MCP 2 during the loading of the $^4\text{He}^*$ MOT, with and without the presence of the ^{87}Rb MOT. The inset shows the ^{87}Rb fluorescence, where after 12 s we start the $^4\text{He}^*$ MOT loading. The drop in ^{87}Rb signal is mainly due to the flux of ground state He atoms. A fluorescence signal of 250 mV corresponds to about 1×10^9 atoms, while the ^{87}Rb loading rate is smaller than in Fig. 4.

−8.4 MHz with respect to the $F = 2 \rightarrow F' = 3$ transition, while the red- and the blue-detuned push beams have a detuning of −8.4 MHz and +17 MHz, respectively. In one of the cooling beams repumper light is mixed in. The differential pumping tube between the 2D- and 3D-MOT sections has a diameter of 2.5 mm and a length of 50 mm, which provides a differential pressure of 1.1×10^{-4} between the 2D- and 3D-MOT sections. With a typical loading rate of 3×10^8 atoms/s we reach 1×10^9 ^{87}Rb atoms in five seconds. Fluorescence signals of the ^{87}Rb 3D-MOT loading for different configurations of the 2D-MOT are shown in Fig. 4.

$^4\text{He}^*$ is loaded from a zero-crossing (spin-flip) Zeeman slower. The detuning of the Zeeman slowing beam is −415 MHz and the power is 9 mW/cm². We obtain $^4\text{He}^*$ 3D-MOT loading rates of about 3×10^8 atoms/s, however, because of strong losses this results in $\sim 5 \times 10^7$ $^4\text{He}^*$ atoms within one second. Ion signals during the $^4\text{He}^*$ MOT loading are shown in Fig. 5, with and without the presence of a ^{87}Rb MOT, indicating a small decrease in $^4\text{He}^*$ final atom number for the two-species MOT compared to single species conditions.

In the experimental sequence we first load ^{87}Rb in the 3D-MOT, while only in the last two seconds we open the *in vacuo* shutter to load $^4\text{He}^*$. During the $^4\text{He}^*$ loading we observe a small decrease of the ^{87}Rb atom number (see inset of Fig. 5), which is mostly due to the flux of ground state He atoms (i.e. this loss is independent on whether the Zeeman slower or $^4\text{He}^*$ 3D-MOT light is on or not).

Afterwards, we compress the 3D-MOT by increasing the gradient to $\alpha = 24$ G/cm in 70 ms, during which we increase the detunings for ^{87}Rb and $^4\text{He}^*$ to −21 MHz

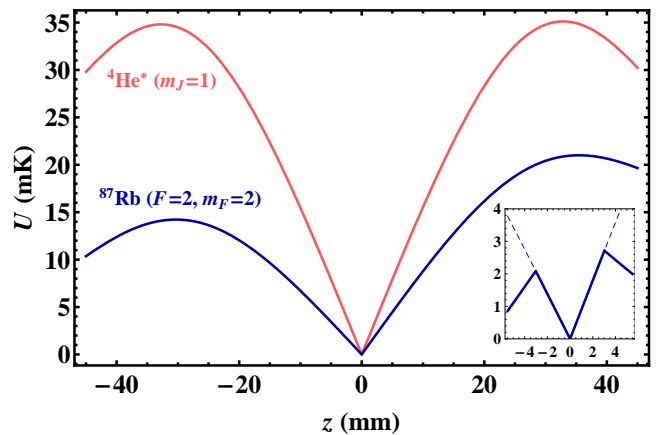


FIG. 6. (Color online) Trapping potentials of $^4\text{He}^*$ (light red) and ^{87}Rb (dark blue) in the quadrupole magnetic trap (QMT) along the z -direction (i.e. along the direction of gravity), for $\alpha = 120$ G/cm. Inset shows the effective potential (solid line) for ^{87}Rb when introducing microwave (MW) radiation 75 MHz above the hyperfine splitting. The noticeable asymmetry in ^{87}Rb potential is due to gravity.

and −44 MHz, respectively. Then we switch off the magnetic field gradient and apply optical molasses on both species for 7 ms. During this optical molasses we ramp the ^{87}Rb detuning from −21 MHz to −29 MHz and lower the intensity of the repumper beam, while for $^4\text{He}^*$ we immediately jump to a detuning of −3.5 MHz. Finally we spin-polarize in 0.5 ms ^{87}Rb to the $F = 2, m_F = 2$ state by optical pumping on the $F = 2 \rightarrow F' = 2$ transition with circular polarized light at a small magnetic field, while at the same time we spin-polarize $^4\text{He}^*$ in the $m_J=1$ state by optical pumping on the $J = 1 \rightarrow J' = 2$ transition with circular polarized light.

D. Ultracold mixture in quadrupole magnetic trap

After the spin-polarizing pulse we ramp the magnetic field gradient within a few ms to $\alpha = 48$ G/cm. After waiting for 100 ms, we ramp to $\alpha = 120$ G/cm in 100 ms. The final trapping potentials are depicted in Fig. 6. More than 95% of the ^{87}Rb atoms are in the $F = 2, m_F = 2$ state, while the $m_J=1$ state of $^4\text{He}^*$ is the only magnetically trappable state. The initial atom numbers for ^{87}Rb and $^4\text{He}^*$ are about 2×10^8 and 5×10^6 , respectively, and their initial temperatures are both about 0.4 mK.

We hold the mixture for a variable time in the magnetic trap, after which we measure the properties of the remaining atoms by absorption imaging (^{87}Rb) and MCP detection ($^4\text{He}^*$). We obtain the time evolution of the atom numbers and temperatures of the two species, in the mixture and under single species conditions. We do not observe any significant effect of $^4\text{He}^*$ on ^{87}Rb , both regarding atom number and temperature, which is mostly

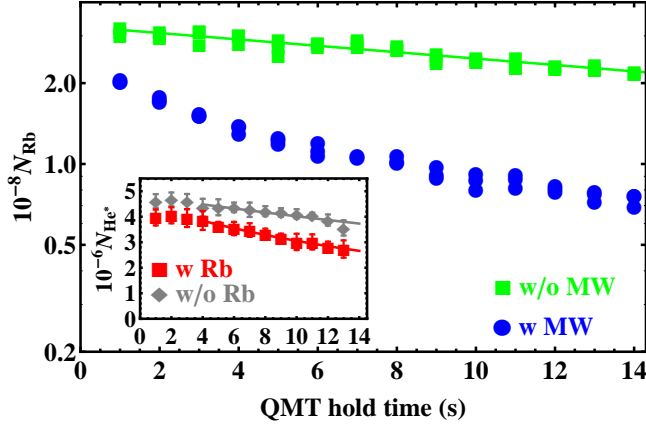


FIG. 7. (Color online) The number of ^{87}Rb atoms as function of hold time in the quadrupole magnetic trap (QMT), with and without microwave (MW). The inset shows the $^4\text{He}^*$ atom number, with (red squares) and without (gray circles) ^{87}Rb and the solid lines are fits of Eq. 10.

explained by the condition $N_{\text{Rb}} \gg N_{\text{He}^*}$.

We apply evaporative cooling on ^{87}Rb by shining in microwave (MW) radiation at 6910 MHz, which is 75 MHz above the hyperfine splitting, leading to an effective trap depth of 2.1 mK (see inset Fig. 6). The time evolution of the ^{87}Rb atom number in the QMT is shown in Fig. 7, with and without MW, while the ^{87}Rb temperature is shown in Fig. 8 (with MW). Without MW we observe an exponential decay of the atom number with a lifetime of 36(2) s, which is due to background collisions. With MW a stronger, non-exponential decay is visible. At a hold time of 4 s we observe from our absorption images that the ^{87}Rb cloud becomes cross-dimensional thermalized, at which point it can be described by a single temperature of 0.33 mK. Afterwards, plain evaporation further reduce the ^{87}Rb temperature to 0.25 mK at 14 s.

The time-evolutions of $^4\text{He}^*$ atom number and temperature are shown in Fig. 7 and Fig. 8, respectively, with and without ^{87}Rb . The initial temperature of $^4\text{He}^*$ is 0.40 mK, after which it increases due to Majorana heating to about 0.50 mK (with ^{87}Rb) or 0.55 mK (without ^{87}Rb) after 13 s. Thus, we observe interspecies thermalization, which however only partly counteracts Majorana heating. In Sec. III E we will determine the interspecies scattering length from this data. For $^4\text{He}^*$ the trap depth is about 30 mK (see Fig. 6), which excludes evaporative cooling and losses.

We also observe a small reduction in the lifetime of $^4\text{He}^*$ in the presence of ^{87}Rb (inset of Fig. 7). The time evolution of the $^4\text{He}^*$ atom number can be described by

$$\frac{d}{dt}N_{\text{He}^*} = -\Gamma N_{\text{He}^*} - L_2 \int n_{\text{He}^*}(\vec{r})n_{\text{Rb}}(\vec{r})d\vec{r}, \quad (10)$$

where Γ is the one-body loss rate due to background collisions and Majorana spin-flips, and L_2 is the total interspecies two-body loss rate coefficient, which in-

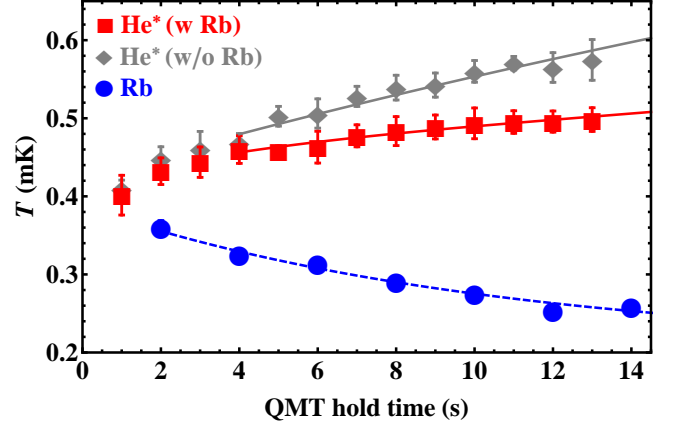


FIG. 8. (Color online) $^4\text{He}^*$ and ^{87}Rb temperatures as function of hold time in the quadrupole magnetic trap (QMT). For $^4\text{He}^*$ the temperature data with ^{87}Rb (red squares) and without ^{87}Rb (gray triangles) are shown, together with the fits of Eq. 9 (red and gray solid lines, respectively), for which only the data from 4 s on is considered. The data points are an average over four experimental runs and the error bars represent the standard deviation. The ^{87}Rb temperature data (blue circles) are fitted by an exponential function of the form $T(t) = (T_0 - T_f)e^{-\beta t} + T_f$ (blue dashed line), which is used as input for solving Eq. 9.

cludes both interspecies PI and spin-relaxation. Intraspecies two-body loss, for which the loss rate coefficient is $2 \times 10^{-14} \text{ cm}^3\text{s}^{-1}$ [30], can be fully neglected.

To extract L_2 , we first fit the data without ^{87}Rb to obtain Γ , after which we fit the data with ^{87}Rb to obtain L_2 . We only give an upper limit of L_2 because the observed reduction in lifetime may also be explained by a few percent of ^{87}Rb atoms in the $F=2$, $m_F=1$ or $F=1$, $m_F=-1$ states, for which PI is not suppressed, or an increase in the Majorana spin-flip loss rate because of the smaller temperature in the presence of ^{87}Rb . We find $L_2^{\text{upper}} = 1.5 \times 10^{-12} \text{ cm}^3\text{s}^{-1}$, which includes the estimated 50% systematic uncertainty in N_{He^*} . L_2^{upper} is three times lower than the reported upper limit of interspecies PI at a temperature of 0.2 mK [6].

E. Determination of the interspecies scattering length from thermalization measurements

To extract the interspecies elastic cross section σ from our data that is displayed in Fig. 8, we first fit the $^4\text{He}^*$ temperature data without ^{87}Rb , which is only described by Majorana heating (Eq. 7), and we find $\chi = 0.09(1)$, similar to 0.14 for ^{23}Na [31] and 0.16 for ^{87}Rb [25]. Then we fit the full solution of Eq. 9 to the $^4\text{He}^*$ temperature data with ^{87}Rb , from which we obtain $\sigma_{\text{exp}} = 14_{-4}^{+6} \times 10^{-14} \text{ cm}^2$. In this analysis we fully take into account the measured time evolution of N_{He^*} , N_{Rb} and T_{Rb} , and propagate their uncertainties (one stan-

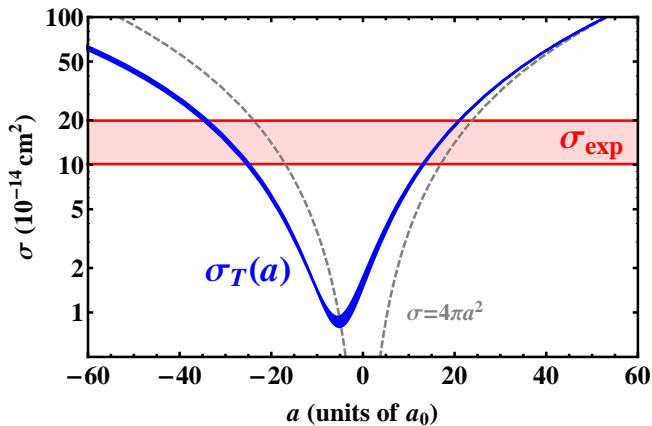


FIG. 9. (Color online) The experimental interspecies elastic cross section σ_{exp} (red horizontal band) for $^4\text{He}^* + ^{87}\text{Rb}$ and the calculated elastic cross section $\sigma_T(a)$, including the contribution from p -wave collisions, for $T=0.45 \rightarrow 0.50$ mK (blue shaded area). The interspecies scattering length is given by the intersections of σ_{exp} and $\sigma_T(a)$, resulting in either $a_{\text{exp}}^- = -29_{-5}^{+5} a_0$ or $a_{\text{exp}}^+ = +17_{-4}^{+4} a_0$. Also the zero-temperature result $\sigma = 4\pi a^2$ is shown (dashed gray line).

dard deviation) to obtain the uncertainty in σ_{exp} . We only fit the data for hold times from 4 s on, at which the ^{87}Rb cloud has become cross-dimensional thermalized. Note that the intraspecies thermalization rate for ^{87}Rb is about $(0.5 \text{ s})^{-1}$ during the whole time evolution, whereas for $^4\text{He}^*$ it decreases from $(0.5 \text{ s})^{-1}$ to $(1.5 \text{ s})^{-1}$.

To relate the temperature dependent elastic cross section to the scattering length we have numerically solved the Schrödinger equation (see Appendix B). It is important to note that such a calculation is only sensitive to the long-range part of the potential, and completely independent of the short-range part obtained from the *ab initio* calculations. The result is depicted in Fig. 9, showing the elastic cross section $\sigma_T(a)$ for the relevant temperature range of 0.45 mK to 0.50 mK as function of scattering length (blue shaded area), which clearly deviates from the zero-temperature limit $\sigma = 4\pi a^2$ (gray dashed line). With our experimental value of σ_{exp} (red horizontal band), we find the scattering length to be either $a_{\text{exp}}^- = -29_{-5}^{+5} a_0$ or $a_{\text{exp}}^+ = +17_{-4}^{+4} a_0$. Note that for the doubly spin-stretched mixture, scattering only occurs in the $^4\Sigma^+$ potential, and the experimentally obtained scattering length is the pure quartet scattering length.

IV. RESULTS AND CONCLUSIONS

The theoretically and experimentally obtained quartet scattering lengths are compared in Table II. One recognizes that a_{exp}^+ is in excellent agreement with the *ab initio* calculations, whereas a_{exp}^- can be fully excluded. In fact, with the bounds of the *ab initio* calculations, we

can conclude that $a_{4+87} = +17_{-4}^{+1} a_0$.

In conclusion, we have performed *ab initio* calculations of the $^4\Sigma^+$ potential for $\text{He}^* + \text{Rb}$, from which we have obtained the interspecies scattering lengths for all four isotope combinations of doubly spin-stretched $\text{He}^* + \text{Rb}$ systems. We have determined experimentally the interspecies elastic cross section for $^4\text{He}^* + ^{87}\text{Rb}$ from thermalization measurements. Our combined theoretical and experimental work provides tight bounds on the interspecies scattering length, which prior to this work was completely unknown. In addition, from our experimental data we obtain an upper limit of the total interspecies two-body loss rate coefficient of $L_2^{\text{upper}} = 1.5 \times 10^{-12} \text{ cm}^3 \text{ s}^{-1}$, which is three times lower than the previous reported upper limit for interspecies PI.

The success of the *ab initio* calculations, being able to quantitatively predict the scattering length for a system containing a heavy, many-electron atom, is linked to the small reduced mass and shallow $^4\Sigma^+$ potential of the $\text{He}^* + \text{Rb}$ system. This leads to a small number of bound states, which reduces the sensitivity of the scattering length to the potential energy curve. Still, to achieve a 1% accuracy of the *ab initio* calculation is a formidable task for a many-electron system. We expect the same level of accuracy for the $^4\Sigma^+$ potentials of any other combination of He^* with an alkali-metal atom.

V. OUTLOOK

The newly obtained knowledge on the scattering lengths is crucial for realizing and exploring quantum degenerate $\text{He}^* + \text{Rb}$ mixtures. For example, the small interspecies scattering lengths will hamper sympathetic cooling of He^* by Rb , and either RF-induced forced evaporation cooling of $^4\text{He}^*$ or sympathetic cooling of $^3\text{He}^*$ with a third species, for which $^4\text{He}^*$ would be an excellent choice [32], is required. Also, on basis of the intra- and interspecies scattering lengths we expect the dual BEC of $^4\text{He}^* + ^{87}\text{Rb}$ to be miscible and stable [33].

The applicability of the ultracold $\text{He}^* + \text{Rb}$ mixture to universal few-body physics, such as the investigation of the Efimov trimer spectrum, crucially depends on the availability and characteristics of interspecies Feshbach resonances. For this purpose close-coupling calculations that include the $^2\Sigma^+$ potential are required. However, *ab initio* calculations of the $^2\Sigma^+$ potential are expected to be less accurate than those for the $^4\Sigma^+$ potential, because the $^2\Sigma^+$ potential is much deeper [34] and supports many more bound states. Experimentally, thermalization measurements in different spin-mixtures might reveal information about the doublet scattering length, however, because Penning ionization is not suppressed, these measurements will be limited by a short lifetime. Therefore we propose to experimentally search for narrow interspecies Feshbach resonances induced by the spin-spin interaction for a mixture prepared in the lower doubly spin-stretched state, which requires a mixture in an opti-

cal dipole trap. The positions of these resonances would reveal the binding energy of the least-bound doublet level, which would provide sufficient information about the $^2\Sigma^+$ potential.

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Appendix A: CCSD(T) calculations

1. Extrapolation towards small internuclear distances

Because the CCSD(T) equations are divergent for the internuclear distances smaller than $r = 8 a_0$, for which the interaction energy is approximately -200 cm^{-1} , we have to extrapolate our results for $r \geq 8 a_0$ towards smaller r in order to describe the repulsive wall up to positive interaction energies. To justify the extrapolation procedure, we have calculated the Hartree-Fock interaction energy (E_{HF}) and correlation contribution to the interaction energy obtained from the coupled cluster doubles (CCD) method (E_{CCD}) at distances close to the inner turning point at about $r = 7.6 a_0$. The results are shown in Fig. 10. Both the Hartree-Fock and correlation contributions behave exponentially [14], which allows extrapolation to distances at which the interaction energy becomes positive.

2. Estimate of accuracy CCSD(T) method

To test the error beyond the CCSD(T) method expansion, we have performed coupled cluster calculations of the interaction energy with singly-, doubly-, triply- and quadruply excited amplitudes (CCSDT and CCSDTQ, respectively) for $r = 9.4 a_0$, using the approach introduced by Kállay [19]. Since the cost of performing these

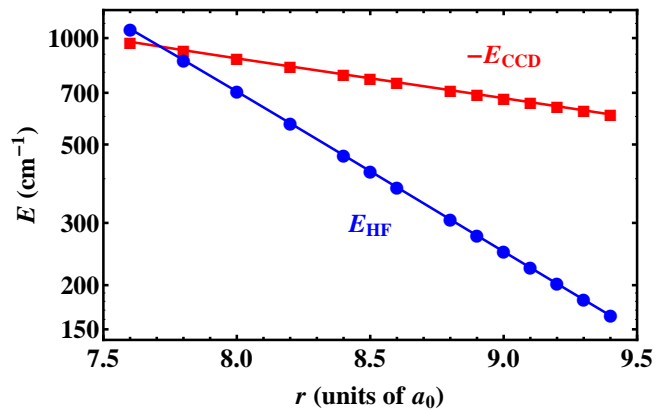


FIG. 10. The contributions of Hartree-Fock (E_{HF}) (blue circles) and correlation effects to the total interaction energy in the CCD method (E_{CCD}) (red squares), both showing an exponential behavior (note the log scale on the Y-axis; lines are a guide to the eye). The two contributions have opposite sign and cancel each other at the inner turning point of the potential energy curve.

calculations is many orders of magnitude higher compared to CCSD(T), we have to restrict ourselves to an 11 valence electrons effective core potential (ECP) and a basis set limited to *spd* orbitals, respectively, and investigated the *difference* with respect to the CCSD(T) interaction energy obtained, *within* the ECP method.

We have found that the inclusion of the full set of triple excitations leads to an increase of the well depth D_e by approximately 3.7 cm^{-1} compared to the CCSD(T) calculations ($D_e = 398.3 \text{ cm}^{-1}$ and 394.6 cm^{-1} , respectively). By further reduction of the basis set (to *sp* orbitals) we have also found that taking into account quadruple excitations has the opposite effect: D_e decreases by 0.3 cm^{-1} compared to the CCSDT calculations. Hence, we can expect that the systematic error due to the incompleteness of the CCSD(T) method should be limited by the difference between CCSDT and CCSD(T) well depths, which in recommended basis set should be proportional to the D_e ratio in the limited basis sets. This leads to a systematic error of $+4.3 \text{ cm}^{-1}$.

Appendix B: Determination of the interspecies scattering length from elastic cross section

In the zero-temperature limit the elastic cross section σ is simply related to the *s*-wave scattering length a via $\sigma = 4\pi a^2$. However, in the temperature range of our measurement, we do not fulfill this limit. Therefore, we have performed numerical calculations on basis of the radial Schrödinger equation to obtain the connection between a and σ at the experimentally relevant temperature range. Here we use a simple Lennard-Jones potential to demonstrate that for this particular purpose

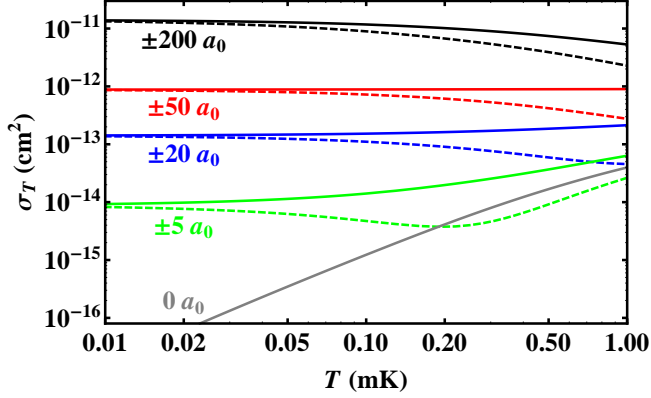


FIG. 11. (Color online) Numerical results of σ_T for $a=\pm 5 a_0$ (green), $\pm 20 a_0$ (blue), $\pm 50 a_0$ (red), $\pm 200 a_0$ (black), where the solid (dashed) lines are representing the positive (negative) values of a , and $a = 0 a_0$ (gray).

only knowledge of the long-range potential is sufficient, and the experimental scattering length determination is completely independent of the *ab initio* calculations of the short-range potential. The Lennard-Jones potential is given by $V_b(r) = -(C_6/r^6) [1 - b/r^6]$, where we can tune the depth of the potential, and therefore a , by parameter b . Note that the less accurate C_6 coefficient of Zhang *et al.* [15] gives the same result than the one obtained in this work, showing that a few percent accuracy is sufficient.

The energy dependent elastic cross section is given by

$\sigma_E = \sum_{l=0}^{\infty} \sigma_E^l$, where $\sigma_E^l = (4\pi/k^2)(2l+1) \sin^2 \delta_l(k)$, δ_l is the l -wave phase shift, $k = \sqrt{2\mu_r E}/\hbar$, μ_r is the reduced mass, and E is the collision energy. The scattering length is defined by $a = -\lim_{k \rightarrow 0} \tan \delta_0(k)/k$. The temperature dependent partial cross section σ_T^l is obtained by taking the Boltzmann average over σ_E^l :

$$\sigma_T^l = \frac{1}{(k_B T)^2} \int_0^{\infty} \sigma_E^l E e^{-E/k_B T} dE \quad (\text{B1})$$

where one has to consider an effective temperature, given by $T = \mu_r(T_{\text{He}^*}/m_{\text{He}^*} + T_{\text{Rb}}/m_{\text{Rb}}) \approx T_{\text{He}^*}$. The total temperature dependent cross section is $\sigma_T = \sum_{l=0}^{\infty} \sigma_T^l$. We find that σ_E^l (and therefore also σ_T) is only dependent on b via a , but different values of b that give the same a , also give the same σ_E^l . This means that our results are independent on the particular choice of model potential, and that σ_T is fully determined by a and C_6 (and the reduced mass μ_r).

In Fig. 11 we show σ_T for several values of a , where the solid (dashed) lines represent the positive (negative) values of a . Because the p -wave centrifugal barrier height is 3.4 mK, in most cases $\sigma_T \approx \sigma_T^0$ for $T \leq 1$ mK. However, the contribution of p -wave collisions, σ_T^1 , can still be significant for small values of $|a|$, for which σ_T^0 itself is very small. Therefore for all calculations we include p -wave collisions, i. e. $\sigma_T = \sigma_T^0 + \sigma_T^1$. In the temperature range of 0.1–1 mK deviations from the zero-temperature cross sections are significant, especially for small $|a|$. In general, for a given $|a|$, the cross section for $a > 0$ is larger than $a < 0$. Finally, we calculate σ_T as a function of a for the experimental relevant temperature range, for which the result is shown in Fig. 9.

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